

Durham Research Online

Deposited in DRO:

18 March 2015

Version of attached file:

Other

Peer-review status of attached file:

Peer-reviewed

Citation for published item:

Lenz, Alexander and Pas, Heinrich and Schalla, Dario (2010) 'Constraints on fourth generation Majorana neutrinos.', Journal of physics : conference series., 259 (1). 012096.

Further information on publisher's website:

<http://dx.doi.org/10.1088/1742-6596/259/1/012096>

Publisher's copyright statement:

© 2010 IOP Publishing Ltd. This is an author-created, un-copyedited version of an article accepted for publication in Journal of physics : conference series. IOP Publishing Ltd is not responsible for any errors or omissions in this version of the manuscript or any version derived from it. The Version of Record is available online at [10.1088/1742-6596/259/1/012096](http://dx.doi.org/10.1088/1742-6596/259/1/012096)

Additional information:

Use policy

The full-text may be used and/or reproduced, and given to third parties in any format or medium, without prior permission or charge, for personal research or study, educational, or not-for-profit purposes provided that:

- a full bibliographic reference is made to the original source
- a [link](#) is made to the metadata record in DRO
- the full-text is not changed in any way

The full-text must not be sold in any format or medium without the formal permission of the copyright holders.

Please consult the [full DRO policy](#) for further details.

Constraints on fourth generation Majorana neutrinos¹

Alexander Lenz, Heinrich Päs and Dario Schalla

Fakultät Physik, Technische Universität Dortmund, 44221 Dortmund, Germany

E-mail: dario.schalla@tu-dortmund.de

Abstract. We investigate the possibility of a fourth sequential generation in the lepton sector. Assuming neutrinos to be Majorana particles and starting from a recent - albeit weak - evidence for a non-zero admixture of a fourth generation neutrino from fits to weak lepton and meson decays we discuss constraints from neutrinoless double beta decay, radiative lepton decay and like-sign di-lepton production at hadron colliders

1. Introduction

The addition of a fourth family of fermions to the known three generations has recently become a popular extension of the standard model (see [1, 2] for a review). Interesting features and benefits of such scenarios include for example:

- weakening of the tension between direct and indirect bounds on the Higgs mass, see e.g. [3, 4, 5],
- a sizeable enhancement of the measure of CP violation, see e.g. [6],
- gauge coupling unification [7],
- new strong dynamic effects due to large Yukawa couplings allowing for dynamical symmetry breaking, see e.g. [8, 9],
- a solution of flavor problems like the observed 3.8σ deviation of the measured value of B_d - B_s mixing from the standard model ($SM3$) prediction [10] which also enhances the dimuon asymmetry of the $SM3$ [11] towards the value measured by the D0 collaboration [12].

In the following we consider the addition of a fourth family

$$\begin{pmatrix} t' \\ b' \end{pmatrix}_L \quad t'_R \quad b'_R \quad \begin{pmatrix} \nu_4 \\ \ell_4 \end{pmatrix}_L \quad \ell_{4R} \quad \nu_{4R}, \quad (1)$$

while the gauge and Higgs sector remains unchanged compared to the $SM3$. Naturally the addition of right-handed neutrinos allows for both Dirac and Majorana mass terms for all four generations of neutrinos. Moreover, tiny neutrino masses are most naturally explained for Majorana neutrinos in a seesaw scheme, so we do not adopt lepton number conservation at this point.

¹ Talk presented by Dario Schalla.

A lower mass bound on SU(2) doublet fourth generation neutrinos can be obtained by the invisible Z -decay width [13]:

$$N_\nu = 2.984 \pm 0.008 \quad (2)$$

constraining the number of light neutrinos with masses $< M_Z/2$ to be three. Consequently

$$m_4 \geq \frac{m_Z}{2} \approx 45.6 \text{ GeV} \quad (3)$$

provides a lower mass bound on fourth generation neutrinos. As the light mass eigenvalues are bounded from above by the Dirac masses – at least in a typical seesaw model, this provides a bound on the Dirac-type mass m_4^D as well. Assuming perturbativity of the fourth generation neutrino Yukawa couplings then constrains Dirac type neutrino masses approximately to the interval

$$45 \text{ GeV} < m_4^D < 1000 \text{ GeV}. \quad (4)$$

Moreover, recent fits to electroweak precision data in a four generation framework ($SM4$) lead to the following constraints on the fourth generation particle spectrum [14]:

$$|m_{t'} - m_{b'}| < 80 \text{ GeV} \quad |m_{l_4} - m_4| < 140 \text{ GeV}. \quad (5)$$

Finally a recent fit to a set of experimental data in the $SM4$ framework has provided some evidence for a non-zero admixture of a fourth generation neutrino, resulting in a PMNS matrix [15]

$$U_{PMNS} = \begin{pmatrix} * & * & * & \begin{smallmatrix} < 0.089 \\ > 0.021 \end{smallmatrix} \\ * & * & * & < 0.029 \\ * & * & * & < 0.085 \\ < 0.115 & < 0.115 & < 0.115 & \begin{smallmatrix} < 0.9998 \\ > 0.9934 \end{smallmatrix} \end{pmatrix}. \quad (6)$$

In the following we use this evidence – as weak as it may be – as a starting point to reconsider bounds on fourth generation neutrino masses.

2. Neutrinoless double beta decay

The most sensitive probe for neutrino Majorana masses is generally neutrinoless double beta decay ($0\nu\beta\beta$). $0\nu\beta\beta$ decay can be realized by the exchange of a Majorana neutrino (see Fig. 1). In the presence of additional heavy neutrino states the usual effective Majorana mass $\langle m_\nu \rangle$ has to be complemented by an effective heavy neutrino mass $\langle m_N \rangle^{-1}$:

$$\langle m_\nu \rangle = \sum_{i=1}^3 U_{ei}^2 m_i \quad \langle m_N \rangle^{-1} = \sum_N U_{eN}^2 m_N^{-1}. \quad (7)$$

The half-life of the decay is then given by [16]

$$\left[T_{1/2}^{0\nu\beta\beta} \right]^{-1} = \left(\frac{\langle m_\nu \rangle}{m_e} \right)^2 C_{mm}^{LL} + \left(\frac{m_p}{\langle m_N \rangle} \right)^2 C_{mm}^{NN} + \left(\frac{\langle m_\nu \rangle}{m_e} \right) \left(\frac{\langle m_p \rangle}{\langle m_N \rangle} \right) C_{mm}^{NL}, \quad (8)$$

where the C_{mm} factors include phase-space factors and nuclear matrix elements [17, 18] and m_p (m_e) the proton (electron) mass. Considering only the heavy neutrino contribution and using the PMNS matrix obtained in Eq. (6) one obtains stringent bounds on the allowed mass range from the current experimental lower half-life bound $T_{1/2}^{\text{Ge}} > 1.57 \cdot 10^{25}$ years [19]. The allowed region is shown in Fig.2.

This leads to the following mass bounds for a single fourth generation Majorana neutrino

$$U_{e4}^{\text{max}} = 0.089 \quad \Rightarrow \quad m_4^{\text{max}} = 6.8 \cdot 10^5 \text{ GeV} \quad (9)$$

$$U_{e4}^{\text{min}} = 0.021 \quad \Rightarrow \quad m_4^{\text{min}} = 3.8 \cdot 10^4 \text{ GeV}, \quad (10)$$

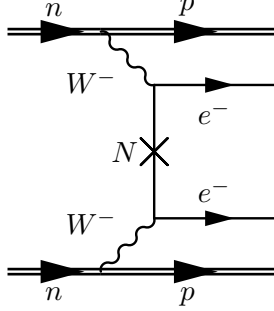


Figure 1. Feynman diagram of $0\nu\beta\beta$ induced by the exchange of a heavy fourth generation Majorana neutrino.

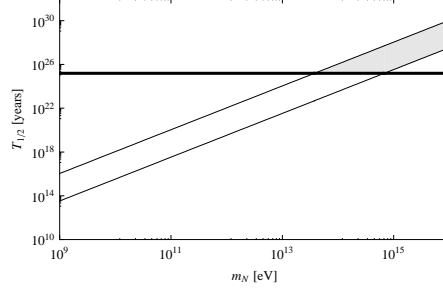


Figure 2. Contribution of heavy neutrino on $0\nu\beta\beta$ half-life (thin lines) within the mixing region given by Eq. (6) and IGEX lower bound (thick line). The gray area indicates the allowed region.

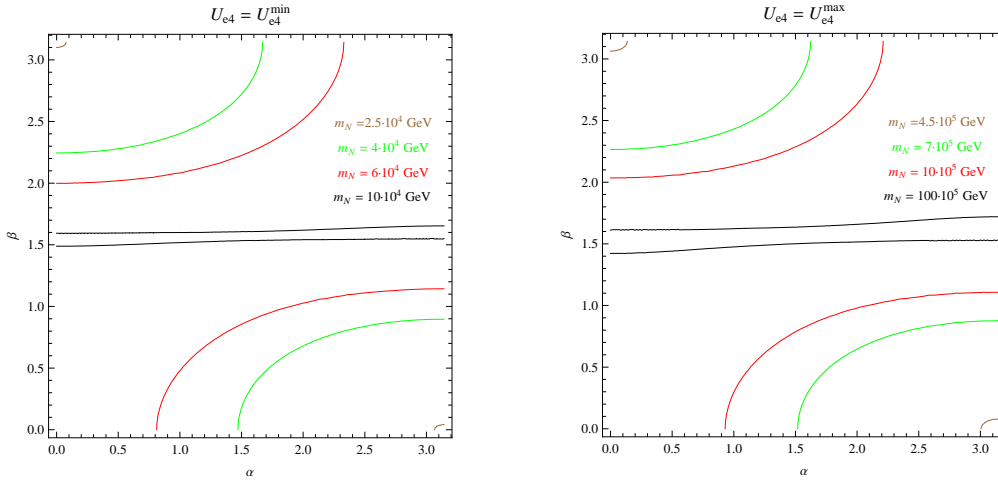


Figure 3. Masses that reproduce the experimental $0\nu\beta\beta$ bound. α is the light and β the heavy neutrino contribution phase.

which are far above the perturbativity constraint of Eq. (4).

Relative phases between light (α) and heavy (β) contributions may cancel each other and thus loosen this bound (see Fig. for several phases α and β). Maximizing the light neutrino contribution by using the largest possible allowed light neutrino masses consistent with the large scale structure of the universe ($\sum m_\nu < 0.66 \text{ eV}$) the mass region of the heavy neutrino can be lowered to

$$2.50 \cdot 10^4 \text{ GeV} < m_4 < 4.49 \cdot 10^5 \text{ GeV}, \quad (11)$$

which remains several orders of magnitude above the desired range (5). In principle there are three different ways to save the possibility of a heavy fourth generation neutrino:

- (i) neutrinos are Dirac particles and therefore $0\nu\beta\beta$ is forbidden, which would come at the cost of seesaw neutrino mass suppression and leptogenesis as a successful way to generate the baryon asymmetry of the universe
- (ii) some other physics beyond the standard model is involved and cancels the heavy neutrino contribution, which would require fine tuning
- (iii) neutrinos are pseudo-Dirac particles.

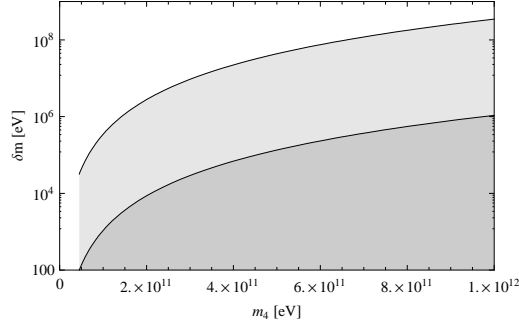


Figure 4. Maximal mass splitting for heavy pseudo-Dirac neutrinos. The upper (lower) curve corresponds to the lower (upper) bound on U_{e4} according to Eq. (6). The marked area represents the allowed parameter space.

In the following we will focus on the latter alternative which may provide useful guidance for future model building.

Pseudo-Dirac neutrinos arise when the Majorana mass is small compared to the Dirac mass. The two resulting mass eigenstates (m_+ , m_-) are nearly degenerate with tiny mass splitting δm and the active and sterile component exhibit practically maximal mixing.

The $0\nu\beta\beta$ half-life of such a neutrino then reads

$$\left[T_{1/2}^{0\nu\beta\beta}\right]^{-1} = \left(\frac{m_p}{\langle m_- \rangle}\right)^2 C_{mm}^{NN} - \left(\frac{m_p}{\langle m_+ \rangle}\right)^2 C_{mm}^{NN}. \quad (12)$$

The allowed mass splittings are shown in Fig. 4 and vary from 32 keV to 350 MeV.

3. Radiative lepton decays

The analysis on neutrino mixing used [15] is dominated by the radiative lepton flavor violating decays of charged leptons. Here we shortly reconsider this bound for the case of pseudo-Dirac neutrinos with masses in the 100 GeV range. The decay width of these processes is given by [20]:

$$\Gamma_{\ell \rightarrow \ell' \gamma} = \frac{1}{2} \frac{G_F^2 m_\ell^5}{(32\pi^2)^2} \alpha |U_{\ell\alpha} U_{\ell'\alpha}|^2 F^2(x), \quad (13)$$

where $F(x)$ is a function of the neutrino masses. It is easy to see that the analysis holds for a pseudo-Dirac neutrino as well. As the fourth generation active and sterile states mix maximally and the masses are close to degenerate $F(x)$ does not change considerably compared to the pure Dirac case. As can be seen from Fig. 5, the decay rate is suppressed by the tiny masses for the first three generations, while the contribution of a fourth heavy generation has to be suppressed due to small mixing.

In the analysis [15] the neutrino mass was fixed to 45 GeV. While a mass dependent study is encouraged, the conclusions of this work will remain unchanged, as the size of the allowed region is anticipated to vary only slightly for different neutrino masses. The decays of the τ lepton do not provide further information as the experimental constraints in this channel are much weaker [21, 22].

4. Like-sign dilepton production

Finally, a process very similar to $0\nu\beta\beta$ is the production of two charged leptons of the same charge at hadron colliders:

$$pp \rightarrow \ell_1^+ \ell_2^+ X. \quad (14)$$

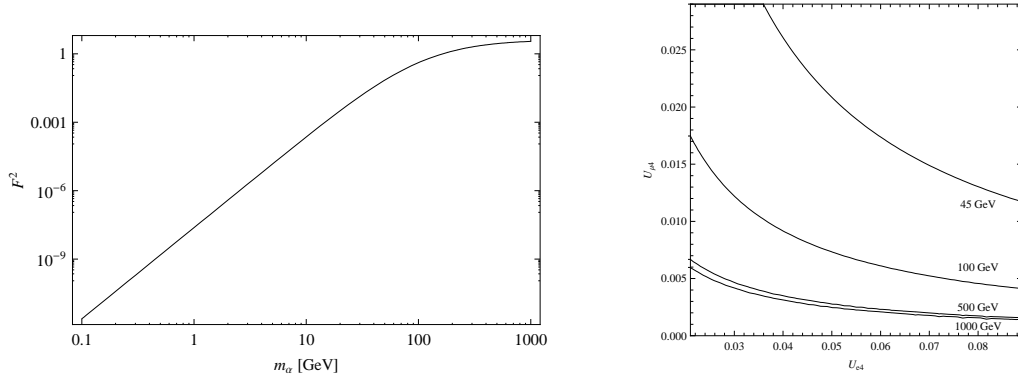


Figure 5. F^2 as a function of the mass of the exchanged neutrino.

Figure 6. Constraint on the $U_{\mu 4} - U_{e 4}$ parameter space obtained from the bound on the branching ratio for $\mu \rightarrow e \gamma$. Allowed is the region to the lower left. The boundaries of the intervalls plotted are given by the allowed values for $m_4 = 45 \text{ GeV}$ according to Eq. (6).

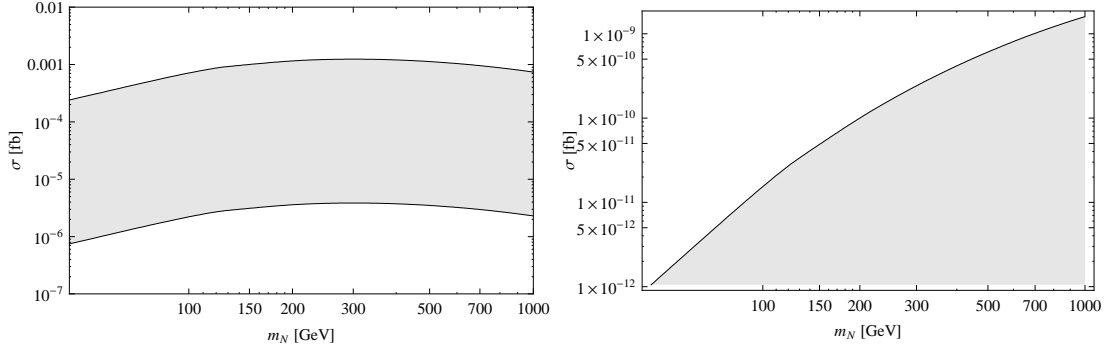


Figure 7. Cross section for like-sign dilepton production by an electroweak scale Majorana neutrino without $0\nu\beta\beta$ constraints.

Figure 8. Cross section for like-sign dilepton production by an electroweak scale pseudo-Dirac neutrino with $0\nu\beta\beta$ constraints.

As shown in Fig. 9 a heavy Majorana neutrino exchange drives the process whose cross section is [23]:

$$\sigma(pp \rightarrow \ell_1^+ \ell_2^+ X) = \frac{G_F^4 m_W^6}{8\pi^5} \left(1 - \frac{1}{2} \delta_{\ell_1 \ell_2}\right) |U_{\ell_1 4} U_{\ell_2 4}|^2 F(E, m_4), \quad (15)$$

where $F(E, m_4)$ is a function of beam energy and neutrino mass. To describe the exchange of a pseudo-Dirac neutrino the cross section has to be modified by introducing a suppression factor (12):

$$\Delta_{pD} \approx 2 \frac{\delta m}{m_4}. \quad (16)$$

Here the mass splitting δm follows from the $0\nu\beta\beta$ constraint and results in a suppression of the mass dependent cross section as shown in Fig. 7 and 8. The resulting cross sections are far to small to be observed at expected LHC luminosities.



Figure 9. Feynman diagrams of like-sign dilepton production

5. Summary

In this note we have revisited bounds on additional Majorana neutrinos, in order to provide a useful guide for fourth generation neutrino model building. We have shown that a fourth generation Majorana neutrino is not yet excluded if it has a mass of several hundred GeV and the Majorana states pair up to form a pseudo-Dirac state. The mixing of such a neutrino is dominantly constrained by the radiative decay of the muon. Due to the pseudo-Dirac nature lepton number violating processes like like-sign dilepton production turn out to be strongly suppressed. Besides being potentially observable in next generation $0\nu\beta\beta$ experiments, the pseudo-Dirac neutrinos could be directly produced at the LHC, as discussed in [24]. In this paper a 5σ discovery reach for heavy neutrino masses up to 100 GeV was advocated with 30 fb^{-1} . While for larger masses the production cross section would decrease, new decay channels open up once the heavy neutrino mass exceeding the Higgs mass, which would require a detailed simulation.

References

- [1] Frampton P H, Hung P Q and Sher M 2000 *Phys. Rept.* **330** 263 (*Preprint hep-ph/9903387*)
- [2] Holdom B *et al.* 2009 *PMC Phys.* **A3** 4 (*Preprint 0904.4698*)
- [3] Novikov V A, Rozanov A N and Vysotsky M I 2010 *Phys. Atom. Nucl.* **73** 636–642 (*Preprint 0904.4570*)
- [4] Kribs G D, Plehn T, Spannowsky M and Tait T M P 2007 *Phys. Rev.* **D76** 075016 (*Preprint 0706.3718*)
- [5] Chanowitz M S 2010 (*Preprint 1007.0043*)
- [6] Hou W S, Mao Y Y and Shen C H 2010 (*Preprint 1003.4361*)
- [7] Hung P Q 1998 *Phys. Rev. Lett.* **80** 3000–3003 (*Preprint hep-ph/9712338*)
- [8] Hung P Q and Xiong C 2009 (*Preprint 0911.3890*)
- [9] Holdom B 2006 *JHEP* **08** 076 (*Preprint hep-ph/0606146*)
- [10] Lenz A *et al.* 2010 (*Preprint 1008.1593*)
- [11] Lenz A and Nierste U 2007 *JHEP* **06** 072 (*Preprint hep-ph/0612167*)
- [12] Abazov V M *et al.* (D0) 2010 (*Preprint 1005.2757*)
- [13] 2006 *Phys. Rept.* **427** 257 (*Preprint hep-ex/0509008*)
- [14] Eberhardt O, Lenz A and Rohrwild J 2010 (*Preprint 1005.3505*)
- [15] Lackner H and Menzel A 2010 (*Preprint 1003.4532*)
- [16] Hirsch M, Klapdor-Kleingrothaus H V and Panella O 1996 *Phys. Lett.* **B374** 7–12 (*Preprint hep-ph/9602306*)
- [17] Hirsch M and Klapdor-Kleingrothaus H V Prepared for International Workshop on Neutrinoless Double Beta Decay and Related Topics, Trento, Italy, 24 Apr - 5 May 1995
- [18] Muto K and Klapdor H V IN *KLAPDOR, H.V. (ED.): NEUTRINOS* 183-237
- [19] Aalseth C E *et al.* (IGEX) 2002 *Phys. Rev.* **D65** 092007 (*Preprint hep-ex/0202026*)
- [20] Cheng T P and Li L F 1980 *Phys. Rev. Lett.* **45** 1908
- [21] Brooks M L *et al.* (MEGA) 1999 *Phys. Rev. Lett.* **83** 1521–1524 (*Preprint hep-ex/9905013*)
- [22] Aubert B *et al.* (BABAR) 2010 *Phys. Rev. Lett.* **104** 021802 (*Preprint 0908.2381*)
- [23] Ali A, Borisov A V and Zamorin N B 2001 *Eur. Phys. J.* **C21** 123–132 (*Preprint hep-ph/0104123*)
- [24] del Aguila F and Aguilar-Saavedra J A 2009 *Phys. Lett.* **B672** 158–165 (*Preprint 0809.2096*)